

The mass-metallicity (MZR) and fundamental metallicity relation (FMR) at $z \sim 1.4$ using VLT-SINFONI near-infrared spectroscopy of zCOSMOS galaxies

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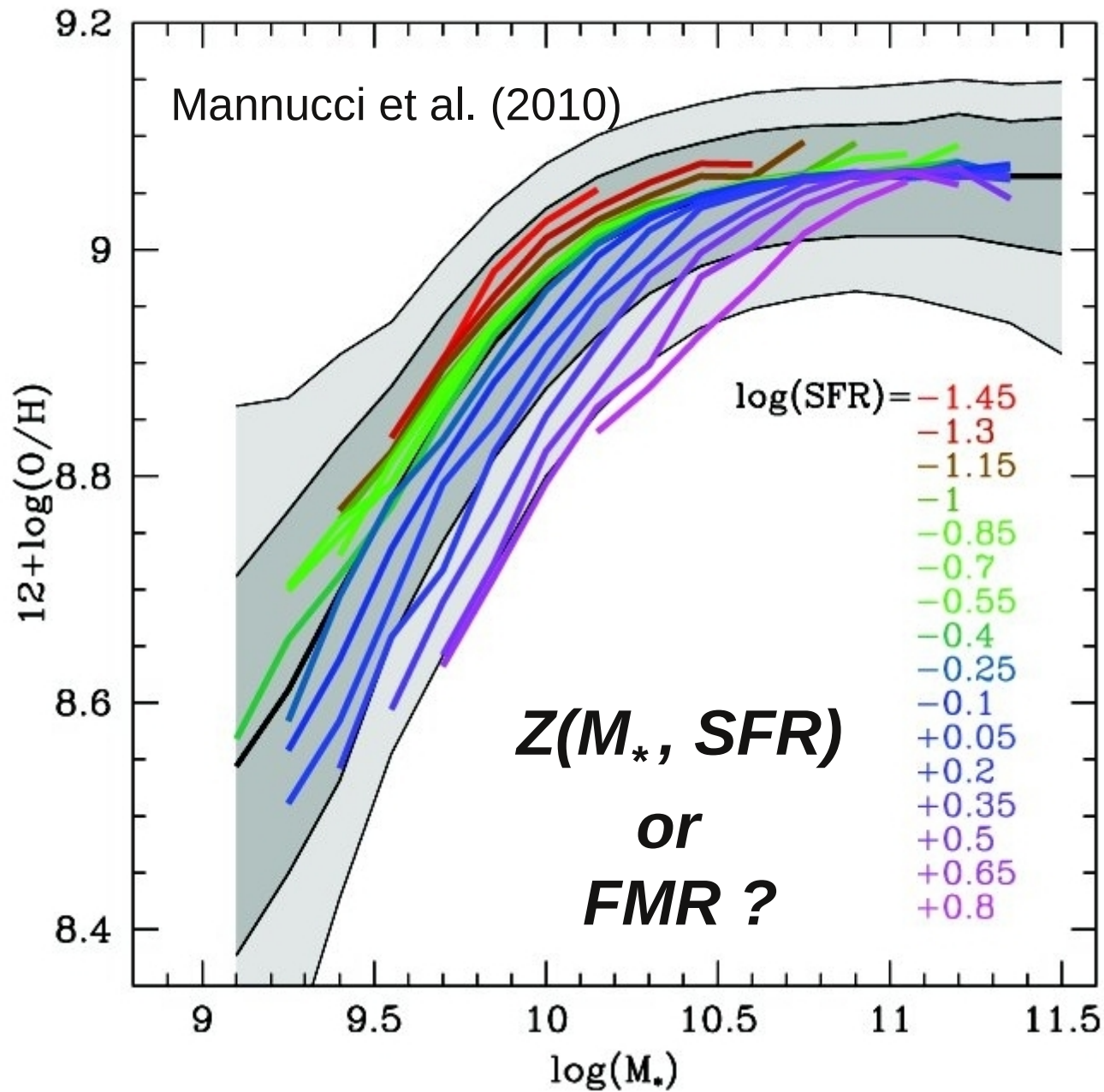


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Empirical dependence: $Z(M_{\text{stellar}}, \text{SFR})$ for SDSS galaxies at $0.07 < z < 0.3$:
at a given mass, galaxies with higher SFRs have lower metallicities

Question: $Z(M, \text{SFR})$ or FMR?

see also discussion in:

Maier, Lilly, Ziegler et al. (2014), ApJ accepted, arXiv:1406:6069

- Is there a dependence of the mass-metallicity relation (MZR) on SFR at $z > 1$, i.e., **does a $Z(M, \text{SFR})$ exist at $z > 1$?**
- **$Z(M, \text{SFR})$ universal (FMR)?**
- **$Z(M, \text{SFR})$ redshift independent or not?**

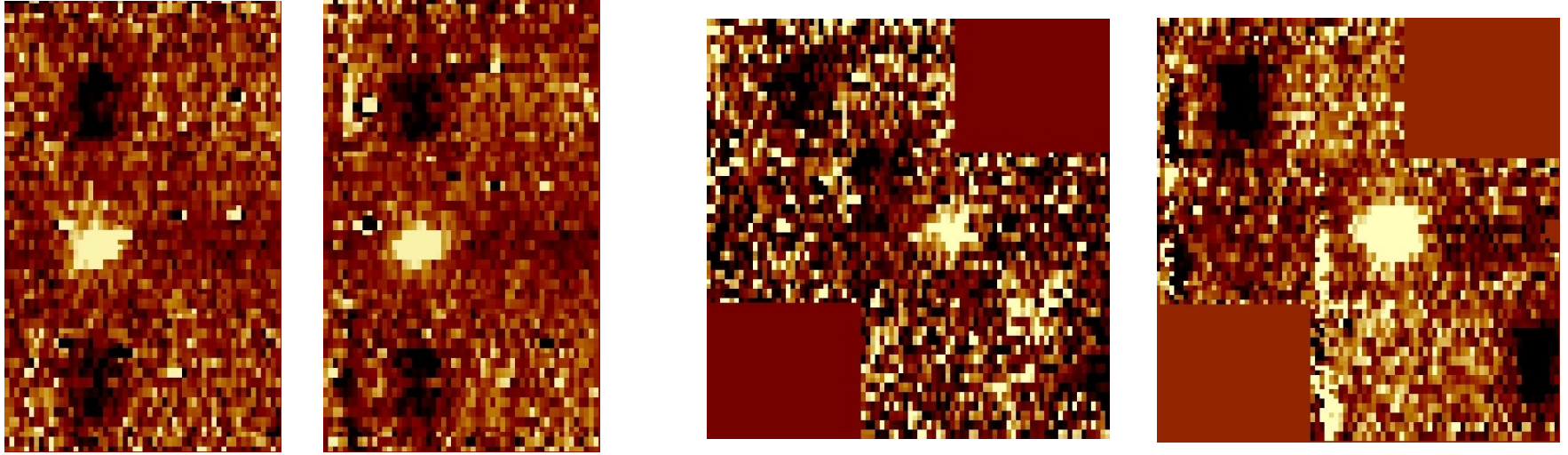
Z: metallicity

M: stellar mass of a galaxy

SFR: star formation rate

FMR: fundamental metallicity relation

Near-infrared spectroscopy with VLT-SINFONI of zCOSMOS galaxies at $z \sim 1.4$



Aim: measure 5 emission line fluxes of $z \sim 1.4$ zCOSMOS galaxies:

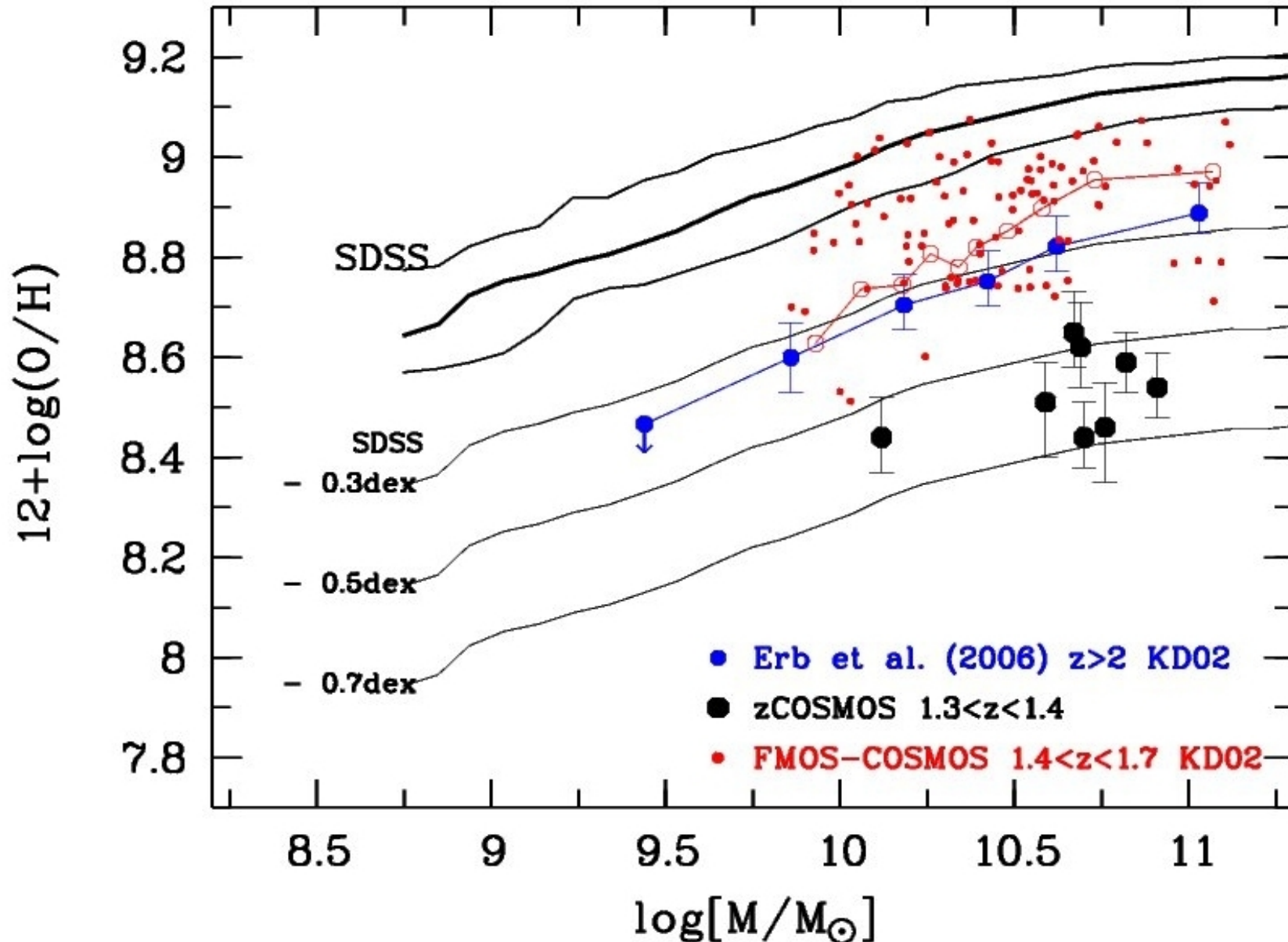
[OII] with VIMOS,

$H\beta$ and [OIII] (SINFONI J-band)

$H\alpha$ and [NII] (SINFONI H-band)

- to measure **reliable metallicities**, which also allows:
- to measure **SFRs** from extinction corrected $H\alpha$
- to identify **Type-2 AGNs** using the BPT diagram
- to study the $Z(M, SFR)$ and FMR

The mass-metallicity relation (MZR) at $z \sim 1.4$



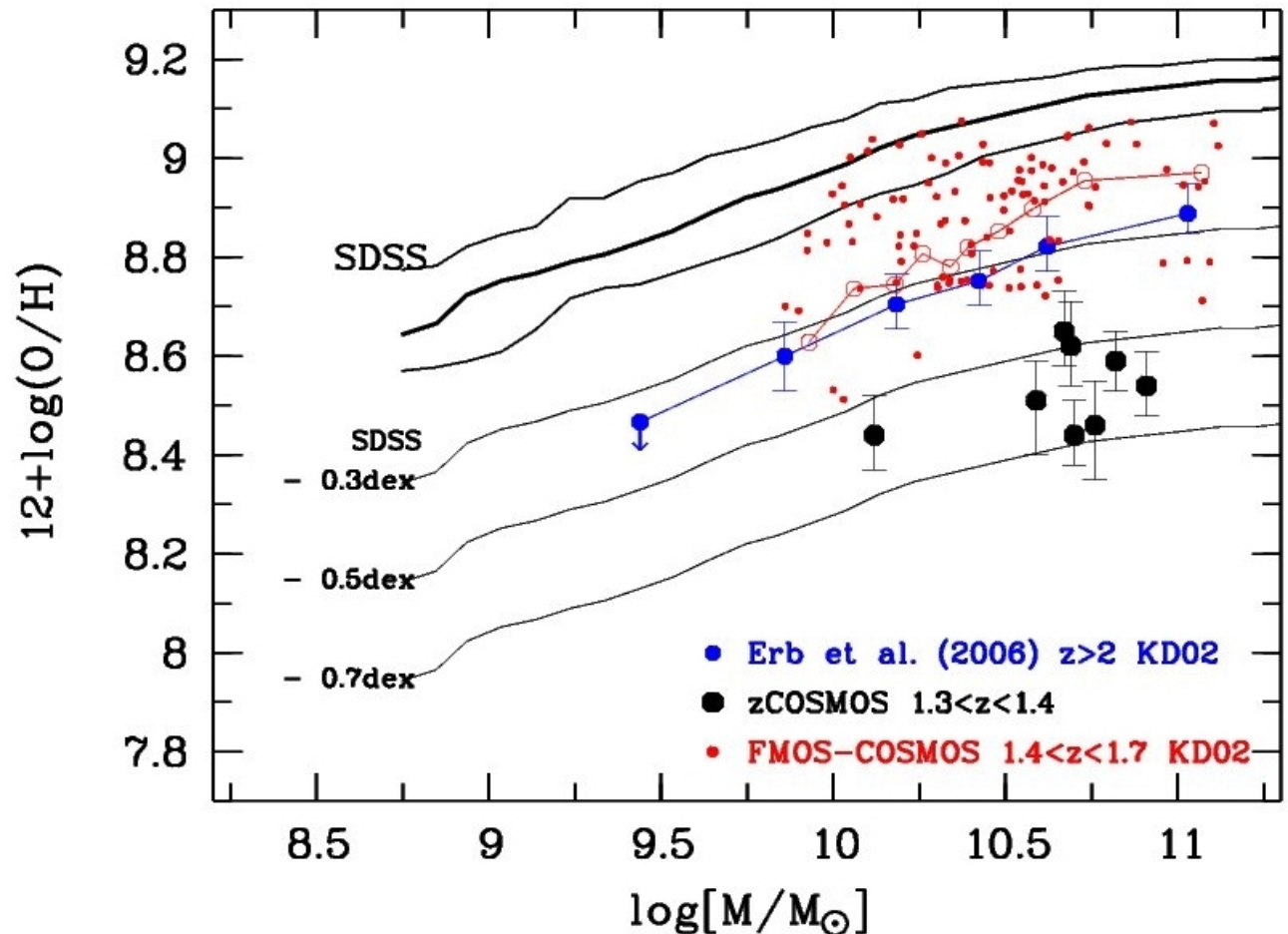
The mass-metallicity relation (MZR) at $z \sim 1.4$

The MZR of $z \sim 1.4$ zCOSMOS galaxies is lower by a factor of 3 to 5 (0.5 - 0.7 dex) than the SDSS relation, while the [NII]/H α -based FMOS-COSMOS O/Hs from Zahid et al. (2014) are lower by up to ~ 2 (0.3dex)

- All O/Hs are transformed to the KD02 (Kewley & Dopita 2002) O/H calibration

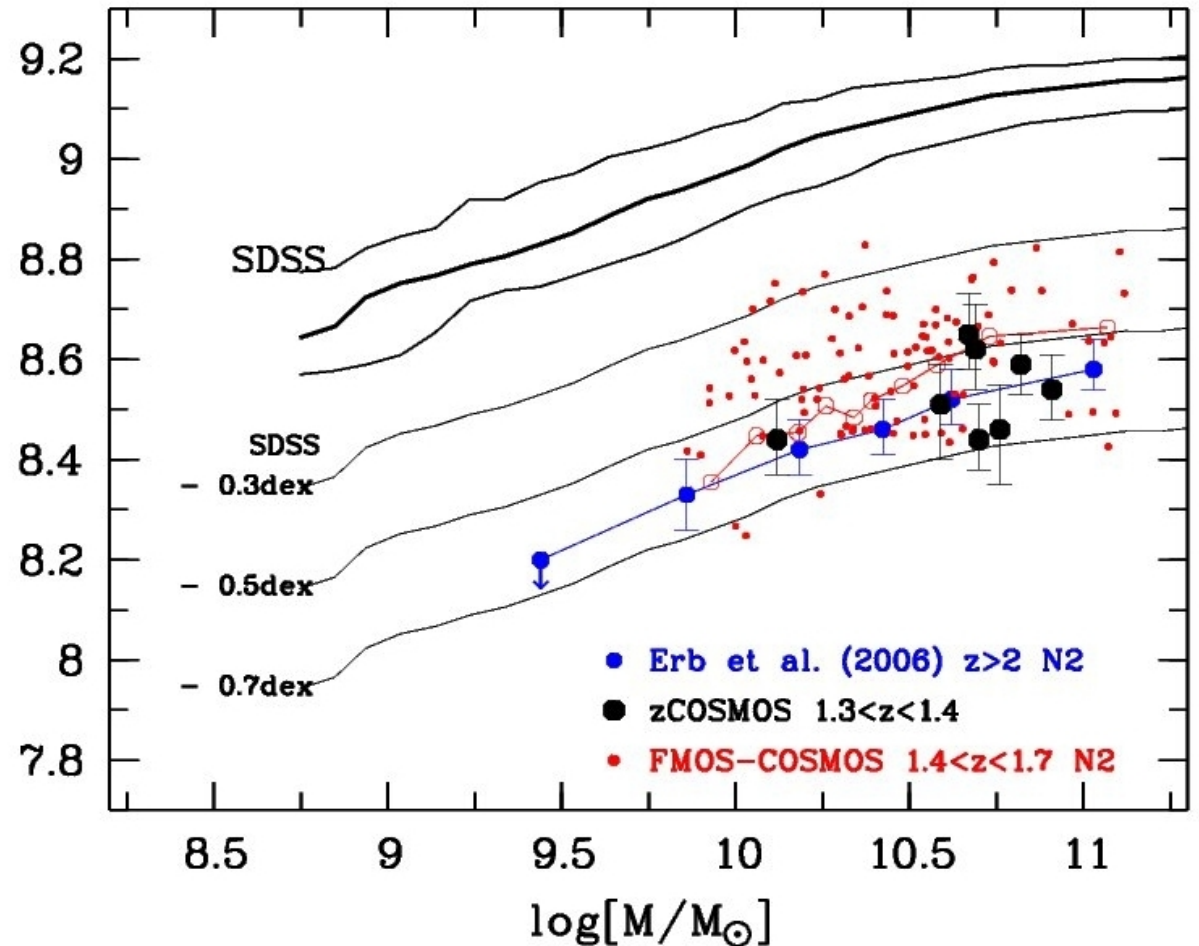
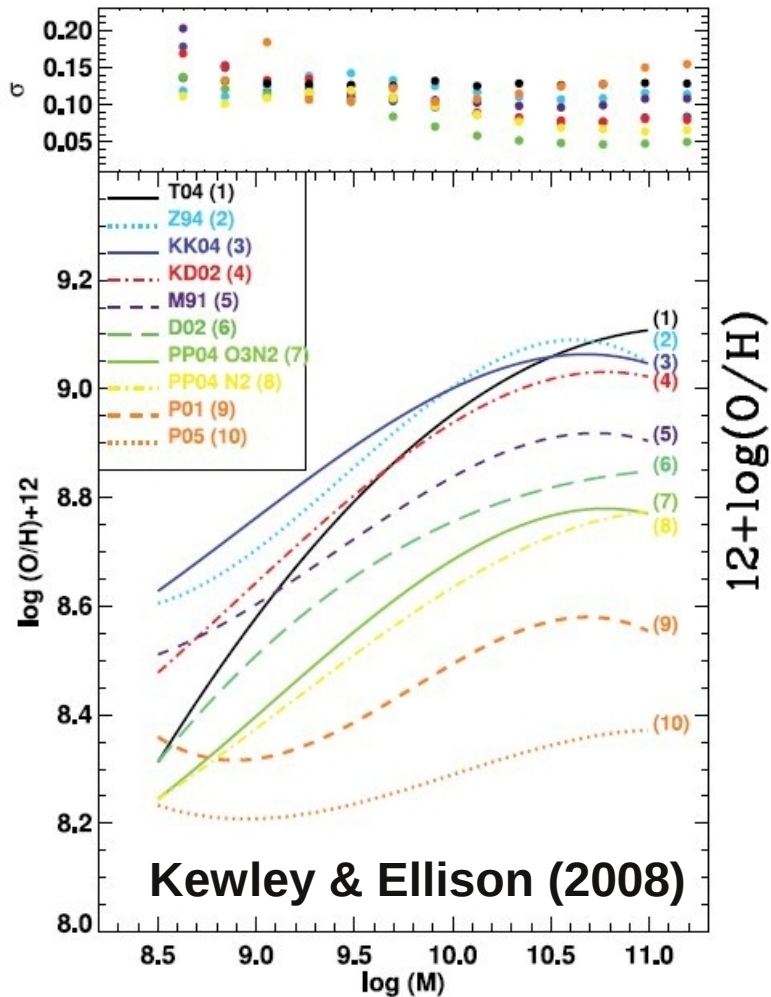
- Newman et al. (2014): **MZR at higher redshifts determined using the N2-method might be 2-3 times too high in O/H**

- Type-2 AGNs contamination in Zahid et al. (2014) sample?

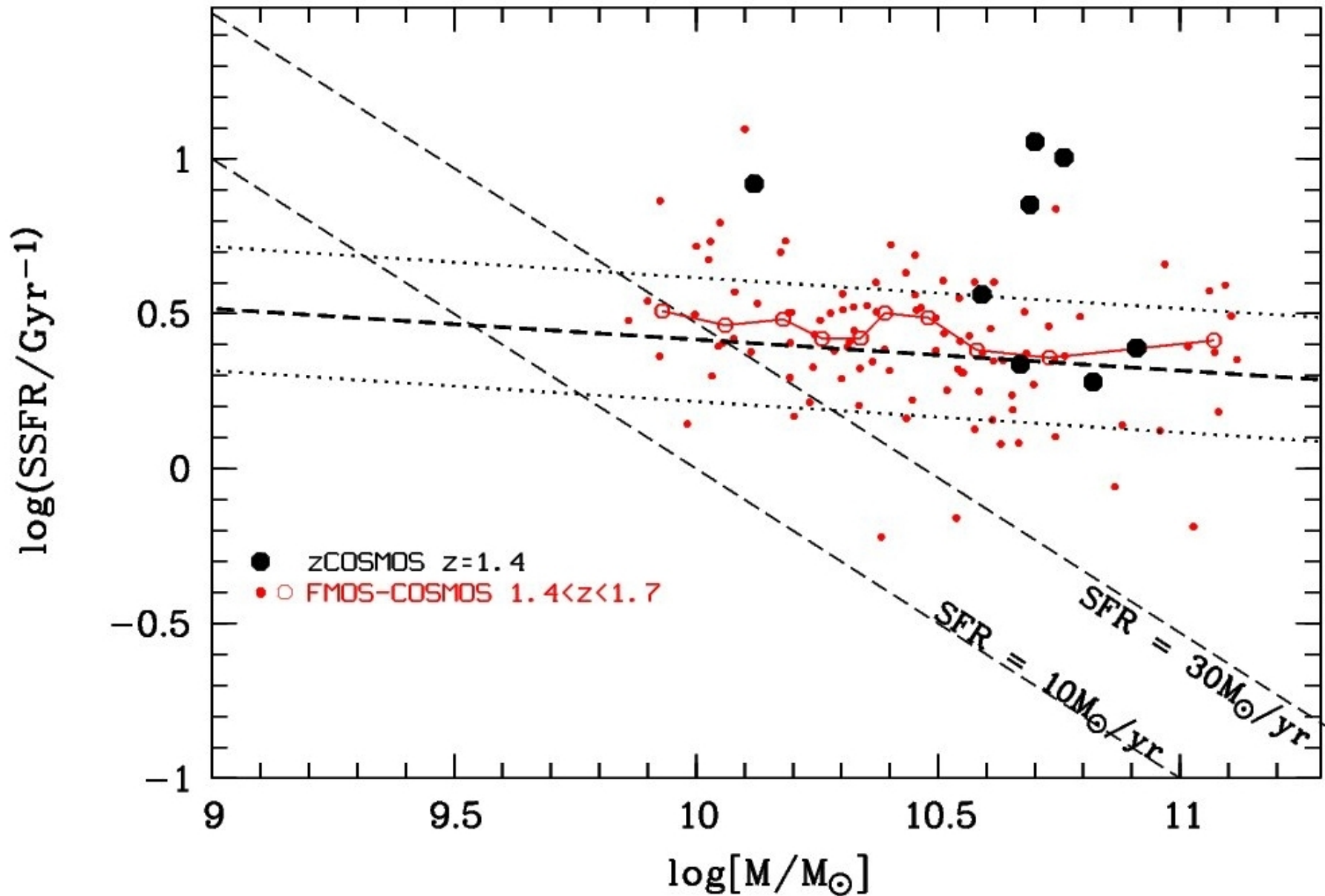


The MZR combining **KD02** and N2 O/H calibrations at $z \sim 1.4$

The mean MZR at $z \sim 1.4$ is lower by a factor of 3-5 (0.5-0.7 dex) than the SDSS relation, when plotting **N2-metallicities** (~ 0.3 dex lower than O/Hs from **KD02**) for $z \sim 1.4$ FMOS-COSMOS galaxies (Zahid et al. 2014)



Sample selection zCOSMOS z~1.4 vs. FMOS-COSMOS

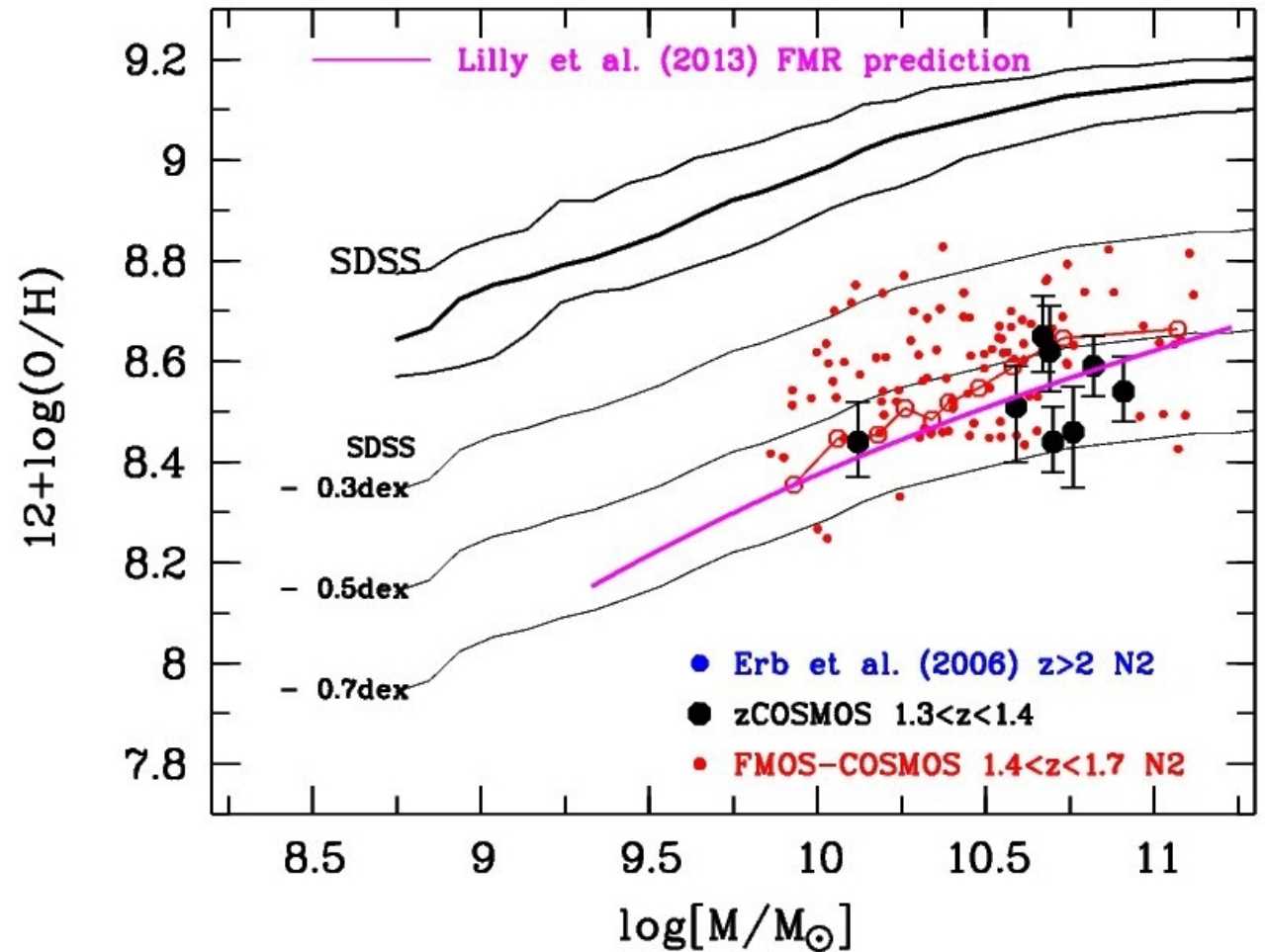
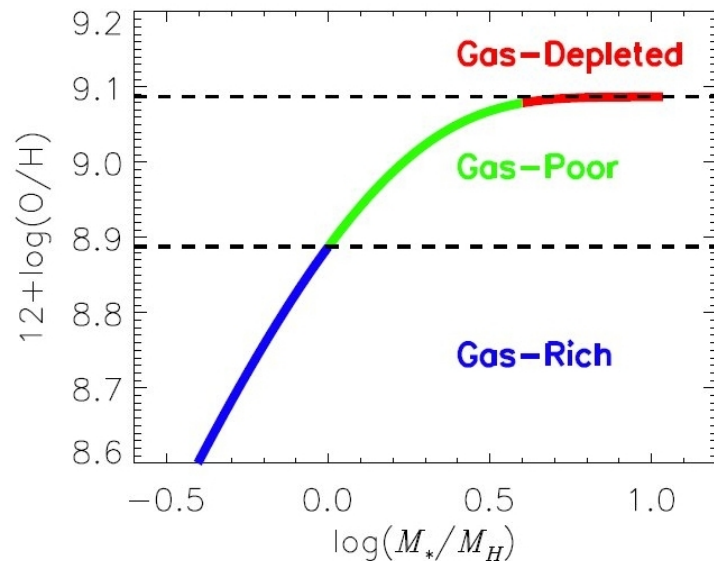


Maier, Lilly, Ziegler et al., in preparation

Z(M, SFR) independent of redshift? FMR?

The $z \sim 1.4$ zCOSMOS and the Zahid et al. (2014) FMOS-COSMOS data are consistent with a non-evolving FMR prediction of Lilly et al. (2013)

Work in progress:
compare the $z \sim 1.4$ observed relation with the model expectations of Zahid et al. (2014b)



$$12 + \log(O/H) = Z_o + \log \left[1 - \exp \left(- \left[\frac{M_*}{M_o} \right]^\gamma \right) \right]$$